The Correlation Between Supernova Fallback and Progenitor's Hydrogen Envelope

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Core Collapse Supernova (CCSN) Mechanism









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Effects of Fallback by Hydrogen Envelope?





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By hydrogen envelope, accretion rate increases

Problem :

H rich

H poor

10⁶

10⁵

- Remnant mass by progenitor hydrogen envelope structure?
- Dependence of explosion energy? ...



How does the hydrogen envelope affect fallback? Motivation

Simulation

Calculation of spherically symmetric 1D fluid in progenitors with "only" different hydrogen envelope

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Initial Condition

Black: type II (Hydrogen: Rich) $Z = 10^{-4}Z_{\odot}, M_{ZAMS} = 18M_{\odot}$ Woosley+02



Simulation Calculation of spherically symmetric 1D fluid in progenitors with "only" different hydrogen envelope



Initial Condition

 $15 \quad 10$

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Simulation Calculation of spherically symmetric 1D fluid in progenitors with "only" different hydrogen envelope

Code: Athena++ (Stone+20) Equations: pure hydro + self-gravity

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{u}) = 0$$

 $\frac{\partial \rho \mathbf{u}}{\partial t} + \nabla \cdot (\rho \mathbf{u} \otimes \mathbf{u} + P^*) = \rho \mathbf{g}$

$$\frac{\partial e}{\partial t} + \nabla \cdot \{ \mathbf{u}(e+p) \} = \rho \mathbf{u} \cdot \mathbf{g}$$

 $p = \epsilon(\gamma - 1), \gamma = 5/3$



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Explosions: Thermal bomb

Inject internal energy (E_{inj}) at $10^7 cm$

⇒Calcurate total energy of ejecta as

explosion energy (E_{exp})

Put $E_{\rm ini}$ to reproduce $E_{\rm exp} \simeq 10^{48-52} {\rm erg}$





- : Hydrogen Rich type II
- : Hydrogen Poor type IIb
- **Before** run in Envelope ⇒**Same**

<u>After</u> run in Envelope



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<u>After</u> run in Envelope















Left panel: Weak fallback!

Right panel: The final remnant mass is ~ $1.8M_{\odot}$, ~ $1.8M_{\odot}$



Result: Explosion energy and the remnant mass Explosion energy (E_{exp}) : the total energy of ejecta $E_{exp} \sim$ total energy of progenitor $+ E_{inj}$



<u>Black: type II (Hydrogen: Rich)</u> Increases $\sim 3M_{\odot}$ by the reverse shock immediately



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 10^{52}



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Black: type II (Hydrogen: Rich) Increases $\sim 3M_{\odot}$ by the reverse shock immediately Transition region: $\sim 2.5 - \sim 6M_{\odot}$ Red: type IIb (Hydrogen: Poor) Shallow than type II The remnant mass is the same at $E_{\rm exp} \gtrsim 4 \times 10^{50} \, {\rm erg}$







Result: Explosion energy and the remnant mass²⁶







Result: Explosion energy and the remnant mass



Parameter search: $M_{\rm ZAMS} = 18,20,24,28 M_{\odot}$ $Z = 10^{-4} Z_{\odot}$

All models have Transition regions!

Transition region: Factor 2-3 of heydrogen envelope's binding energy





Summary

Simulation

Calculation of spherically symmetric 1D fluid in progenitors with **"only"** different hydrogen envelope

Result

- Reverse shock makes Transition region
- Transition region mass range is $2.5M_{\odot}$ to $6M_{\odot}$ at $M_{ZAMS} = 18M_{\odot}, Z = 10^{-4}Z_{\odot}$
- Transition region: $2 3 \times E_{\text{grav,Hyd}}$

Future Task • What is the factor of 2-3?

- The effect of changing $M_{\rm ZAMS}$ and Z ?
- Multi-dimensional effect?



 $E_{\rm exp}$ [erg]





How to make inner region



Self-gravity of thin matter pointmass $\nabla p_{\rm m}(r) + G\rho_{\rm m}(r) \left(\frac{m_{\rm m}(r)}{r^2} + M_{\rm pt} \nabla \phi_{\rm pt}'(r, r_{\rm s}) \right) = 0,$

 $\nabla p_{\rm m}(r_{\rm s}) = \nabla p(r_{\rm s}),$

 $M_{\rm pt} + m_{\rm m}(r_{\rm s}) = m(r_{\rm s}).$

 $0 \leq q < 1;$



Result : Explosion energy and M_C Explosion energy (E_{ex}) : the total energy of ejecta





Result : Normalized Explosion energy and M_C

$$x = \frac{E_{\text{ex}}}{E_{\text{g,H}}}, y = \frac{M_C - M_{\text{pt}}}{M_* - M_{\text{pt}}}$$

 $y = 0: M_{\rm C} = M_{\rm pt}$ no accretion

 $y = 1: M_* = M_C$ Full collapse

cutoff : $x \sim 2 - 3$ $x = 10^{-1}$: all lines have $y \simeq 0.8$ $x \gtrsim 3$: only u18 is different ?



 ${\mathcal X}$















