#### Illuminating Scalar Dark Matter Co-Scattering with Monophoton Signatures

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Based on

arXiv: 2508.06040

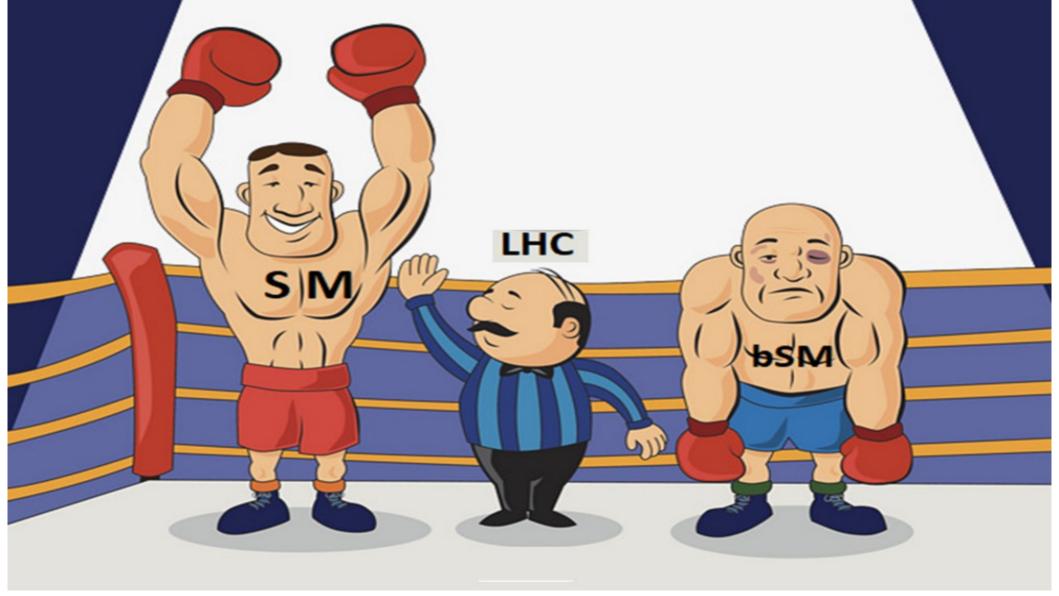
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Scalars 2025– Warsaw, Poland

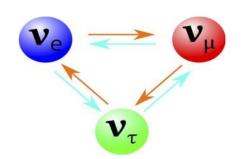
24 September 2025



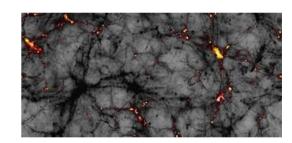


#### Do we have a good reason to go Beyond the Standard Model?

• SM fails to explain neutrino mass and mixings.



SM doesn't have DM candidate.



SM fails to explain observed baryon asymmetry.



# Who can be a DM?

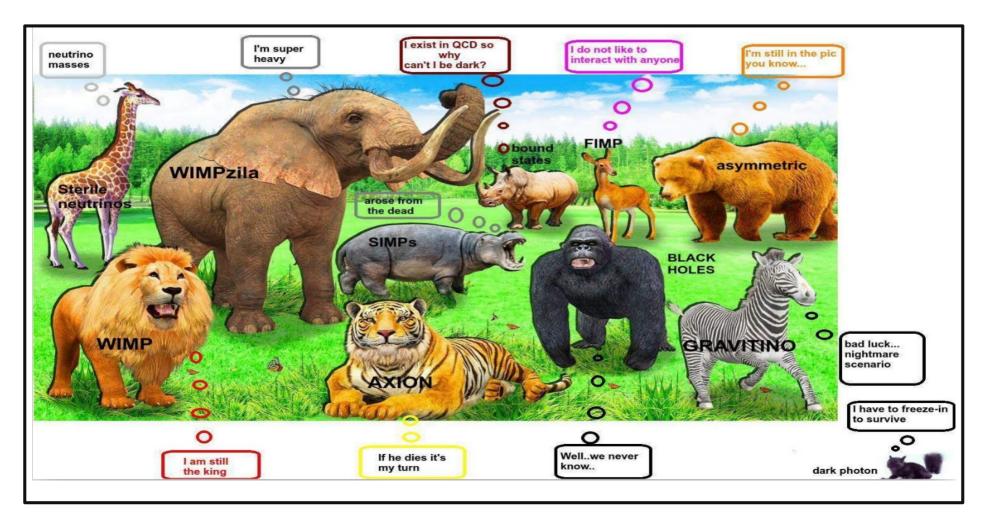
- Should be massive
- Should be electrically neutral
- Should be present in early universe
- Should be stable or at least with half life greater than the age of the universe

  Need a

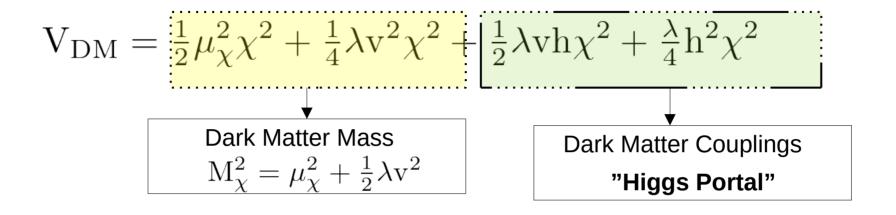
symmetry

Singlet Scalar Singlet Fermion Scalar in triplet reprise Fermion in triplet reprise ...and many more

# **Zoo of Dark Matter Candidates**



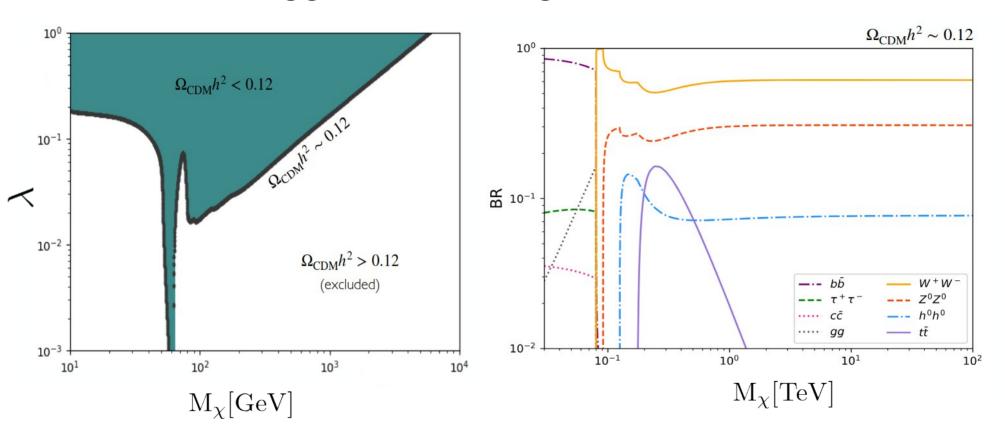
## <u> Higgs Portal : Singlet Scalar DM</u>



Simplest extension of the Standard Model...

- $\rightarrow$  dark matter: real scalar singlet (stable due to  $\mathcal{Z}_2$  imposed symmetry)
- phenomenology (at the tree-level) governed by only two parameters
- One coupling (to Higgs) drives all DM observables DM relic, Direct Detection, Indirect Detection.

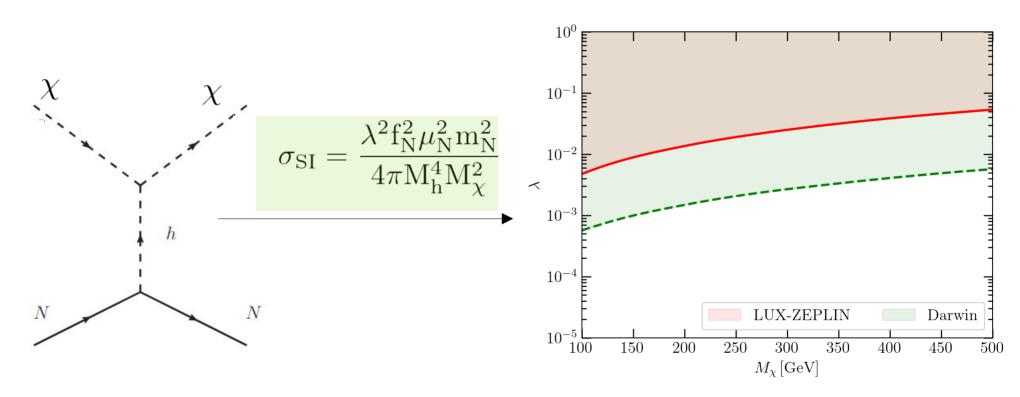
# **Higgs Portal: Singlet Scalar DM**



Dark matter annihilation into: gauge bosons, Higgs bosons, quarks, leptons.

## **Higgs Portal: Singlet Scalar DM**

Higgs portal interactions give spin-independent nuclear scattering via t-channel Higgs exchange.



Direct detection limits imply that the Higgs-portal coupling must be suppressed.

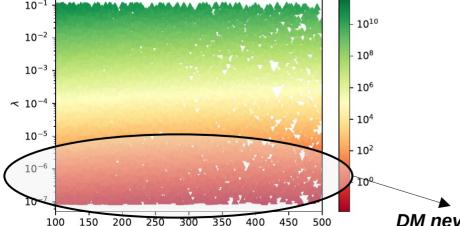
## **Higgs Portal: Singlet Scalar DM**



Requires large Higgs Portal Couplings



Requires Suppressed Higgs Portal Couplings



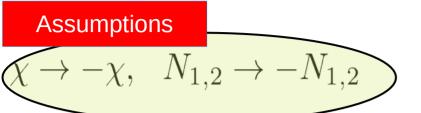
 $M_{\chi}$  [GeV]

 $(\Gamma/H)_{T=M_\chi}$ 

• If Singlet Scalar is **viable WIMP DM**, we need alternate production mechanism to realize the observed DM relic density.

DM never thermalizes, it behaves as non-thermal particle

## <u>Singlet Scalar DM + dimension-5 Operators</u>



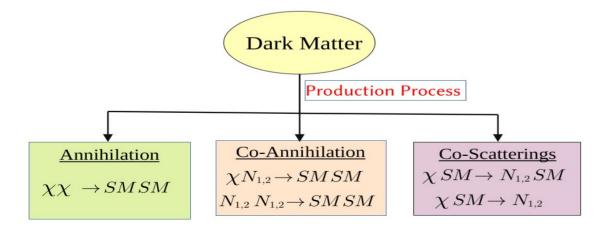
Tree level neutrino mass: forbidden by  $Z_2$  symmetry



$$\mathcal{L}_{eff} \supset 0 \Phi^{\dagger} \Phi \chi^{2} + \frac{c_{5}}{\Lambda} (\overline{L^{c}} \tilde{\Phi}) (\tilde{\Phi}^{\dagger} L) + \frac{Y}{\Lambda} \overline{L} \tilde{\Phi} N \chi + \frac{c_{3}}{\Lambda} \overline{N^{c}} \sigma_{\mu\nu} N B^{\mu\nu}.$$

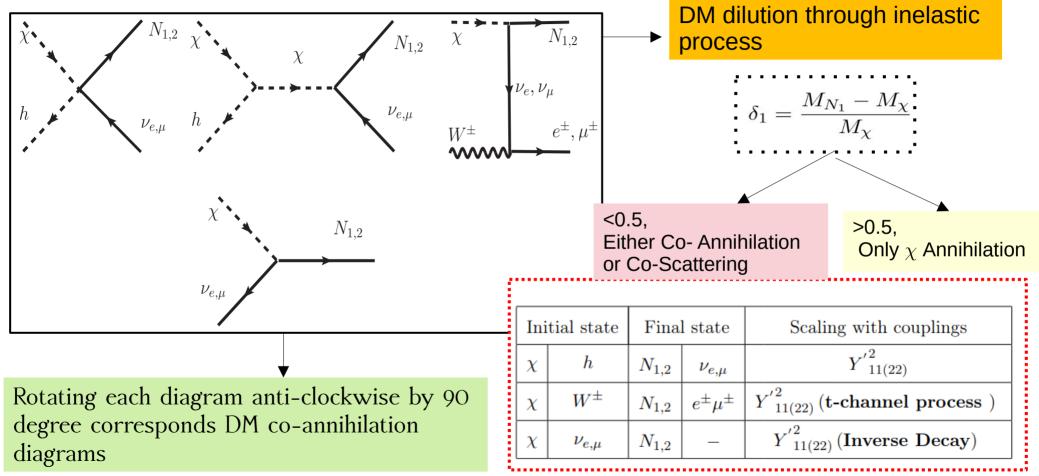
Suppressed due DD constraint

Introduces additional DM dilution processes

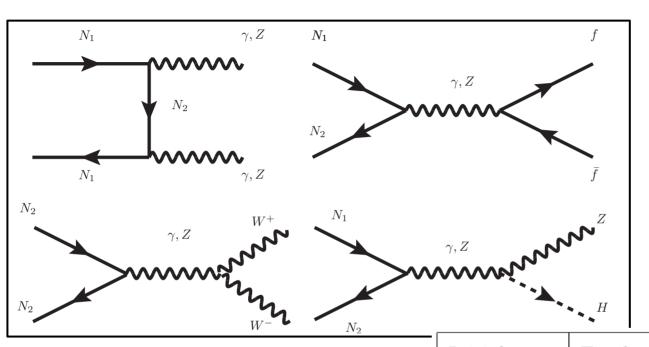


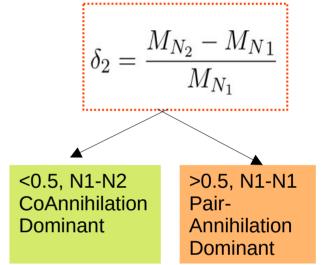
- $\rightarrow \chi$  and  $N_{1,2}$  may or may not be in equillibrium with each other.
- $\Omega_{\chi}h^2$  is set either through coannihilation or co-scattering.

## Singlet Scalar DM + dimension-5 Operators



## **Singlet Scalar DM + dimension-5 Operators**





Dominant co-annhilation process  $\blacktriangleleft$ Contribute less than 1% due to cancellation b/w the Z and  $\gamma$  exchange diagram

	Initial state		Final state		Scaling with couplings
	$N_{1,2}$	$N_{1,2}$	$\gamma, Z$	$\gamma, Z$	${c_3^{\prime}}^4 ( ext{t- channel process}\ )$
	$N_1$	$N_2$	f	$\bar{f}$	${c_3^{'}}^2 ( ext{s-channel process})$
_	$N_1$	$N_2$	$W^+$	$W^-$	$c_3^{'2}  ( ext{s-channel process}  )$
	$N_1$	$N_2$	Z	H	$c_3^{'2}( ext{s-channel process})$

## **Coscattering equations (conversion-driven freeze-out)**

- if DM is very weakly coupled to the SM, DM self-annihilation is negligible
- in the following, 0: SM, 1: N (=N1 + N2), 2: Dark Matter DM

$$\frac{dY_1}{dx} = -\frac{1}{x^2} \frac{s(M_\chi)}{\tilde{H}(M_\chi)} \left[ \langle \sigma_{1100} v \rangle (Y_1^2 - Y_1^{eq2}) - \underbrace{\Gamma_{2 \to 1}}_{s} \left( Y_2 - Y_1 \frac{Y_2^{eq}}{Y_1^{eq}} \right) \right],$$
 Inelastic Processes 
$$\frac{dY_2}{dx} = -\frac{1}{x^2} \frac{s(M_\chi)}{\tilde{H}(M_\chi)} \left[ \frac{\Gamma_{2 \to 1}}{s} \left( Y_2 - Y_1 \frac{Y_2^{eq}}{Y_1^{eq}} \right) \right].$$
 SM SM

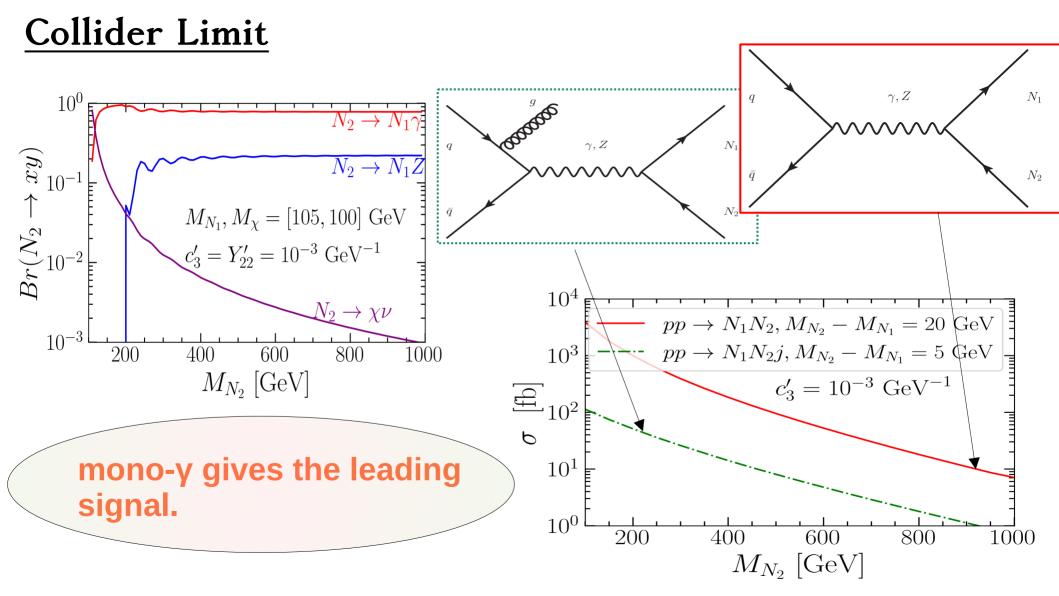
 $Y', \lambda \sim \mathcal{O}(10^{-6}-10^{-10})$  DM pair annihilation, co-annihilation, and exchange Process becomes negligible

# Coscattering fraction

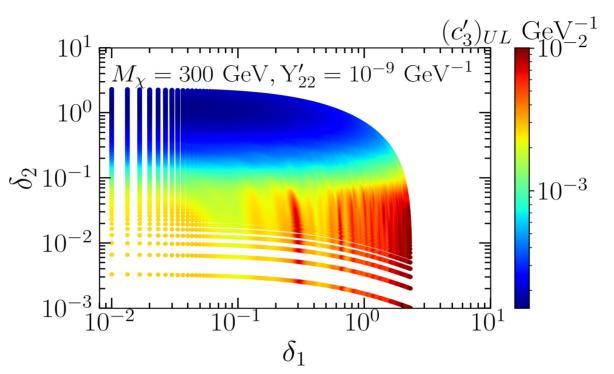
- When Y' and  $\delta_1$  is small, co-scattering keeps  $\chi$  coupled to the N(=N1+N2).
- without coscattering, DM freezes out very early ⇒ too high relic density
- to quantify when coscattering is necessary to keep  $\chi$  coupled to the N(=N1+N2).

$$\Delta_{\chi}^{1} \equiv 1 - \frac{\Omega h^{2}(\text{Single})}{\Omega h^{2}(\text{Coupled})}.$$

- → if co-annihilation dominant  $\Rightarrow \Delta_{\chi}^{1} = 0 \, (Y > 10^{-7})$
- $\rightarrow$  if co-scattering dominant  $\Rightarrow \Delta_{\chi}^{1} = 1 (Y < 10^{-7})$

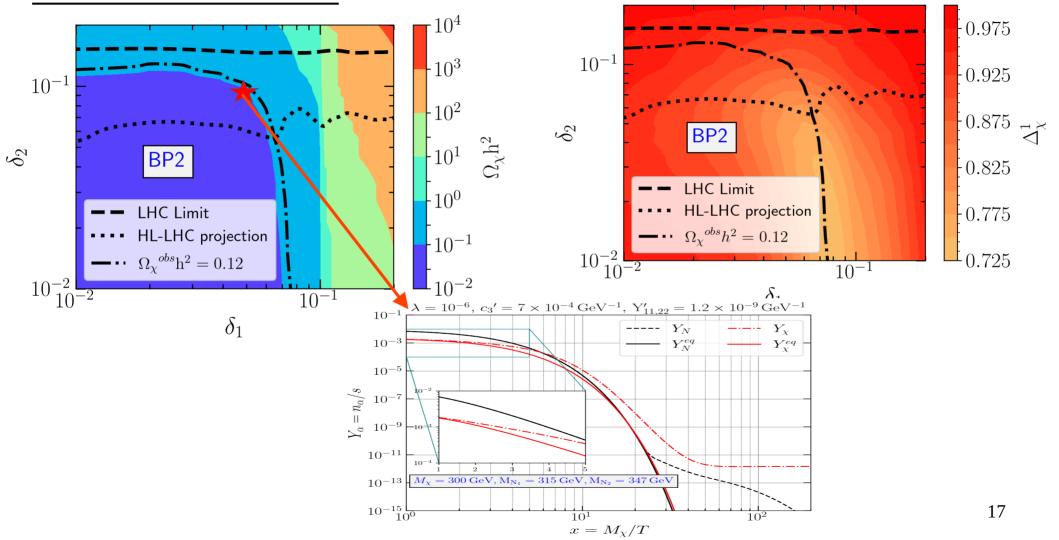


## Numerical results



- "Dark matter + energetic photon in ATLAS (arXiv: 2011.05259): parameter space constraints
- Large  $\delta_2$  leads to energetic photons. Hence, stringent constraints.

#### Numerical results



# Summary

#### **Features of co-scattering dark matter:**

- small coupling to visible matter
- compressed dark sector
- freeze-out works for a wide range of energies

Singlet Scalar DM + dim-5 operators: consistent with DD, ID & collider bounds

Viable parameter space can be probed at HL-LHC

